### Minisymposium on Analysis and Representation of Large Data Sets

# Visualization on large data volumes in Physics

Seeing is believing?



Madrid, 15 February 2012

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(special thanks to E.Martinez, F.Castejon, L.Cabellos)

Outlook

### Examples:

- Event Displays in Particle Physics
- The Cosmic Microwave Background Maps
- Visualization of Trajectories in a Fusion device

### What else:

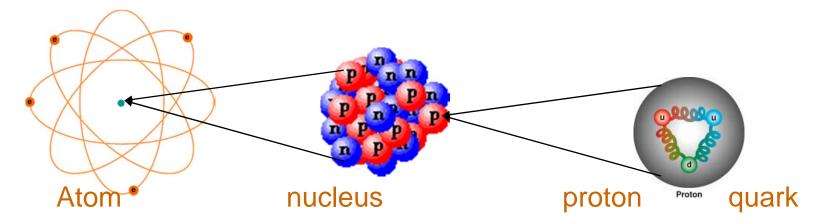
- Dataflow hardware and software
- Complex systems



### Searches in Particle Physics

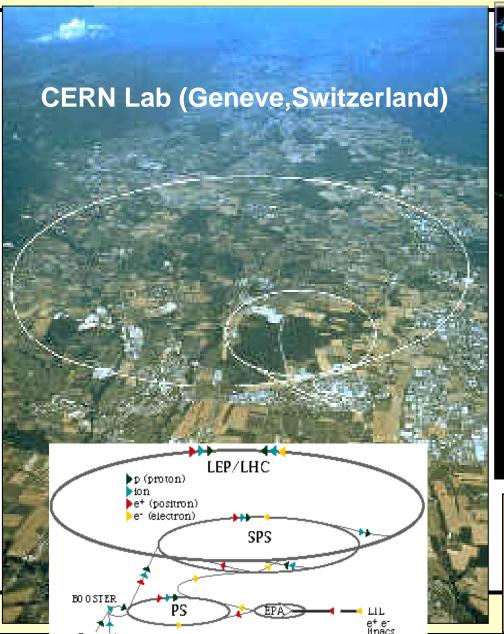
Particle physics: studying the basic constituents of all matter around!

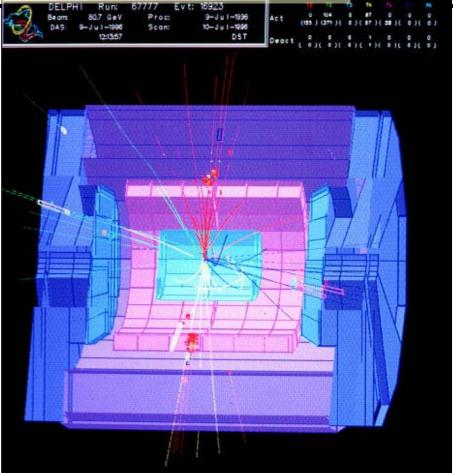




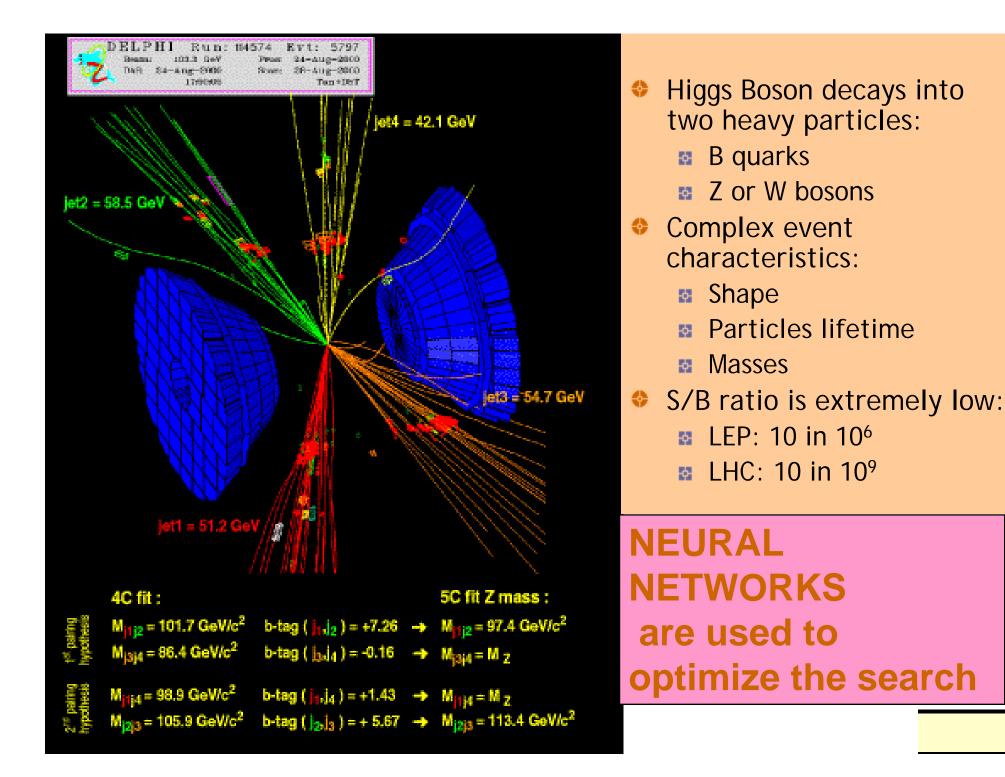
The origin of the *"mass"* of all particles is linked to a fundamental particle predicted but not yet discovered: **the Higgs boson** 

### Accelerators and detectors





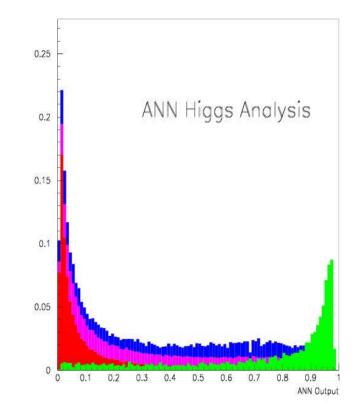
LEP (ended in 2000) Large Electron Positron Collider E<sub>cm</sub>= 200 GeV, e+e- collisions Search for Higgs up to M=115 GeV



# How did we select that "collision"?

#### ANN: example of architecture 16-10-10-1

- 16 input variables
- 2 hidden layers with 10 nodes each
- 1 output layer, 1=signal, 0=background
- Trained on MC sample
  - Biggs generated at a given mass value
  - All types of Background
  - 10x real data statistics
- Applied on real collected data to order in S/B the candidates to Higgs boson
- Training process:
  - Minimize classification "error"
  - Iterative process
  - No clear "best strategy"
- Computing Intensive: hours to days for each "try"



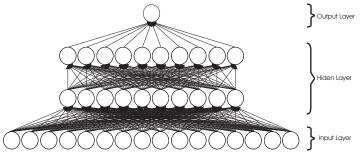
# Distributed Analysis

#### Distributed Configuration:

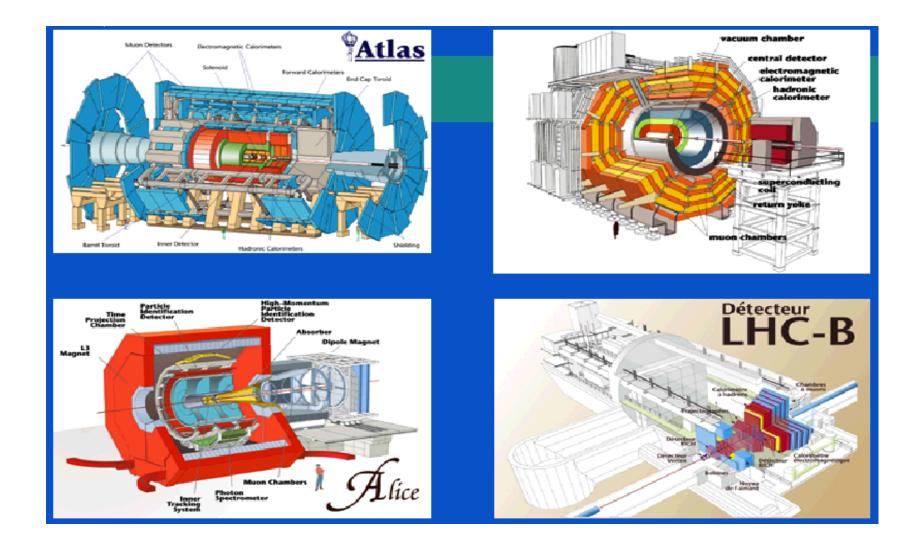
- Master node and N worker nodes.
  - Scan to filter events & select variables
  - ResultSet in XML, split according to N (number of slave nodes)

#### Training procedure:

- Master reads input parameters and sets the initial weights to random values.
- The training data is distributed to the workers.
- At each step:
  - The master sends the weights to the workers.
  - The workers compute the error and the gradient and return them to the master.
- This training procedure has been implemented using MPI and adapting the MLP-fit package.
- Conditions:
  - train an ANN with 650.000 simulated realistic LEP events, 20000 of them corresponding to signal events.
  - Use a 16-10-10-1 architecture (270 weights)
  - Need 1000 epochs training.
  - Similar sized samples for the test.
  - BFGS training method.



### LHC Experiments





### The GRID success!

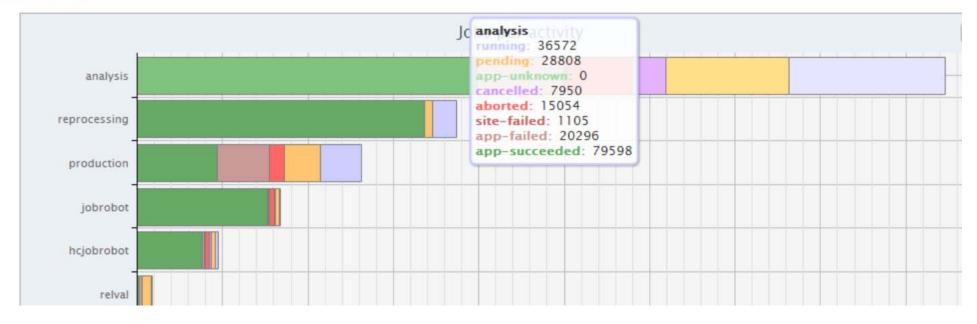
PRODUCTION Normalised CPU time (kSI2K) by REGION and VO Asia Pacific CERN 1000mT NGI\_AEGIS NGI\_ARMGRID 900m NGI BG NGI\_BY 800m NGI\_CH NGI\_CYGRID 700m+ NGI\_CZ NGI\_DE 600m NGI\_FRANCE NGI\_GRNET 500m+ NGI\_HR 400m NGI\_HU NGI\_IBERGRID 300m NGI\_IE NGI\_IL 200m NGI\_IT NGI\_NDGF 100m NGI\_NL NGI\_PL 01 NGI\_RO ems. allice NGI\_SI NGI\_SK NGI\_TR\_SGA 'EGI View': PRODUCTION / normcpu / 2011:3-2012:2 / REGION-VO / Lhc (x) / ACCBAR-LIN / i NGI UK 2012-02-13 02:01

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# Processing...

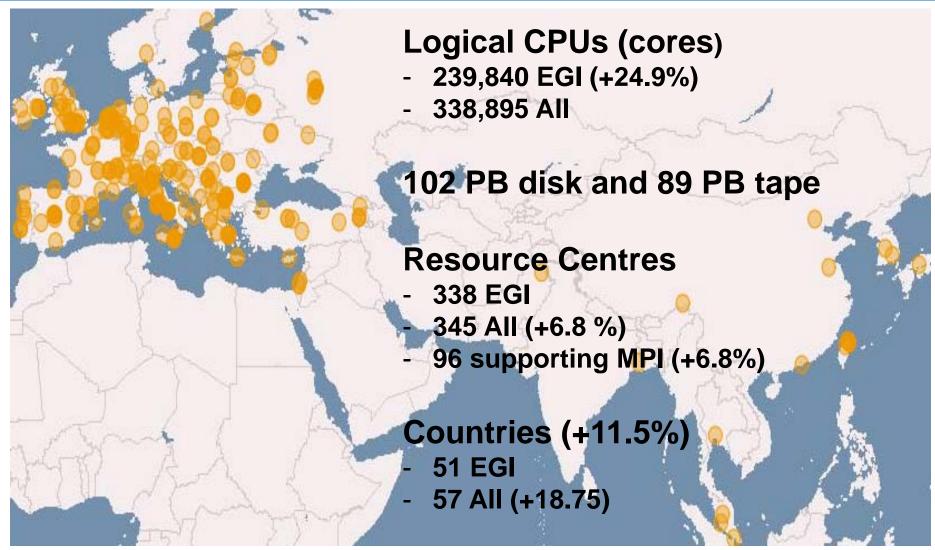
#### Main Jobs Chart





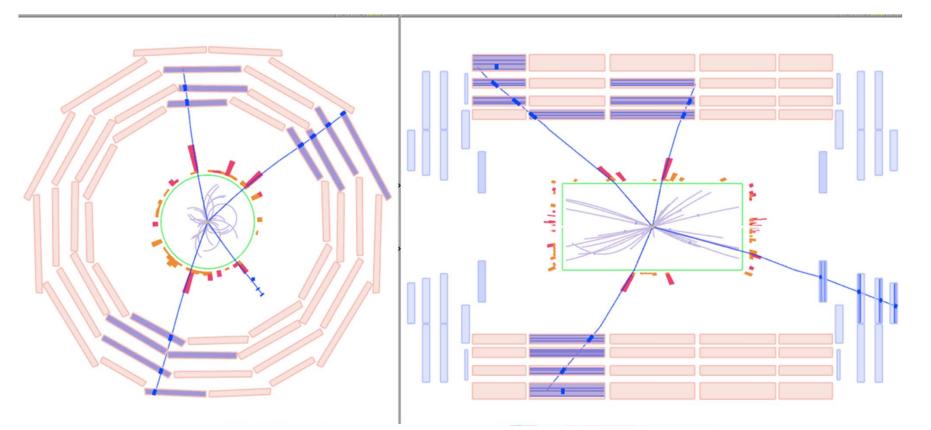


# European Grid Infrastructure (2011)



30/05/2011 EGI-InSPIRE RI-261323 Project Presentation - May 2011

### LHC collisions: a ZZ event



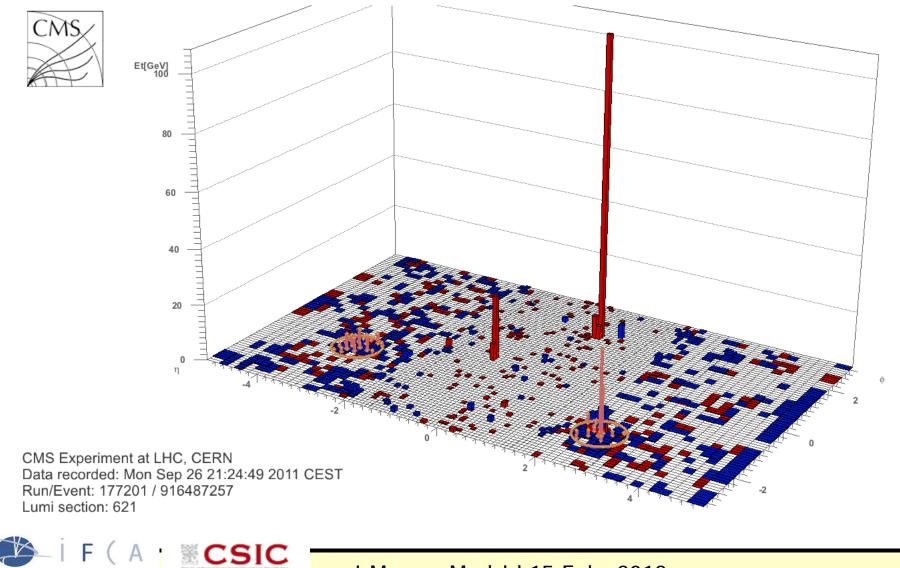
#### **Invariant Masses**

 $\mu_0 + \mu_1$ : 92.15 GeV (total(Z)  $p_T$  26.5 GeV,  $\phi$  -3.03),  $\mu_2 + \mu_3$ : 92.24 GeV (total(Z)  $p_T$  29.4 GeV,  $\phi$  +.06),  $\mu_0 + \mu_2$ : 70.12 GeV (total  $p_T$  27 GeV),  $\mu_3 + \mu_1$ : 83.1 GeV (total  $p_T$  26.1 GeV).



LHCC open n Invariant Mass of 4µ: 201 GeV

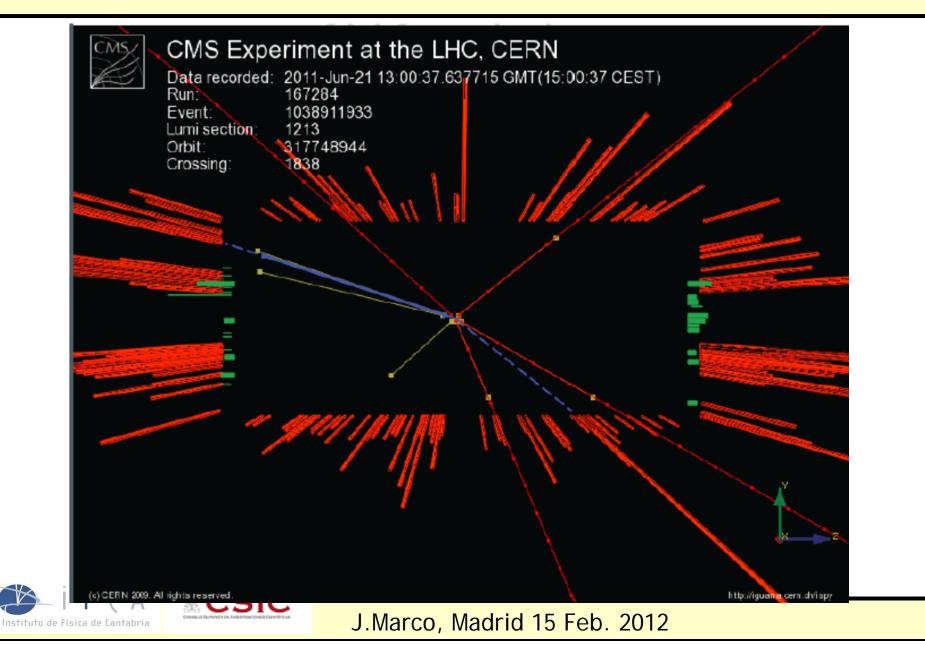
# Higgs candidate decaying into $\gamma\gamma$



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# Higgs into 4 leptons





CMS Experiment at LHC, CERN Data recorded: Mon Nov 8 11:30:53 2010 CEST Run/Event: 150431 / 630470 Lumi section: 173

> One of the first HI central collisions

CMS report LHCC open session, T. Camporesi, 17 November, 2010

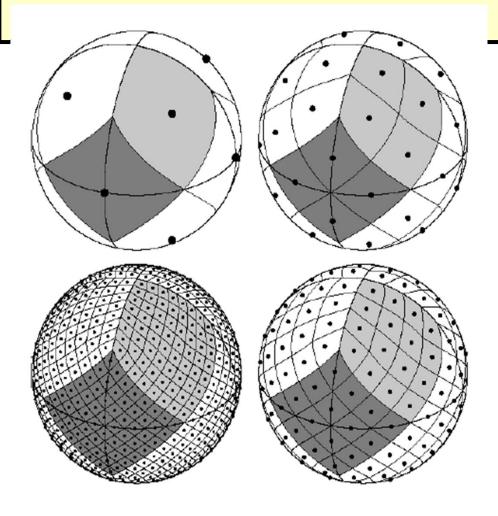
# Symmetries in the Sky

### An example: spherical topology

- Very relevant in Astronomy
- But also in geophysics and in atomic and nuclear physics
- "(There is) no known point set analog to uniform sampling in Euclidean space"
- CMB Sky images
  - Topic: Cosmic Microwave Background (CMB) anisotropies
  - Data: multifrequency/high resolution temperature and polarization microwave sky observation
    - NASA WMAP mission
    - ESA PLANCK mission
  - O(200 Mbytes) images



# Discretization

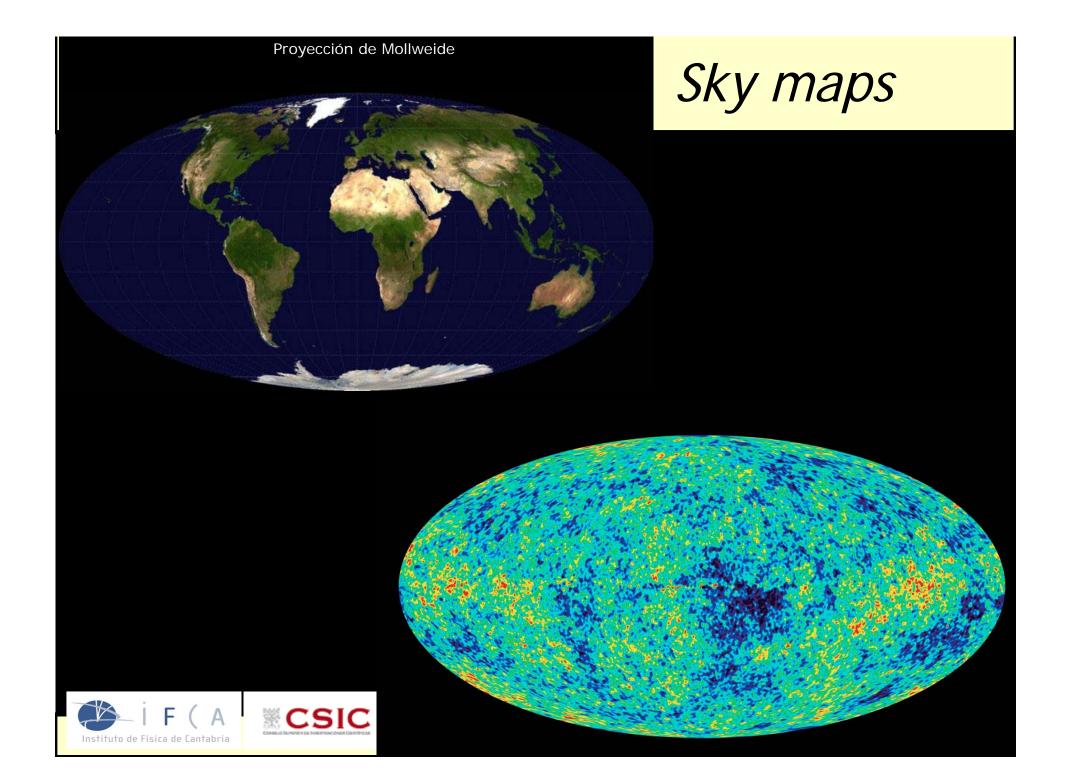


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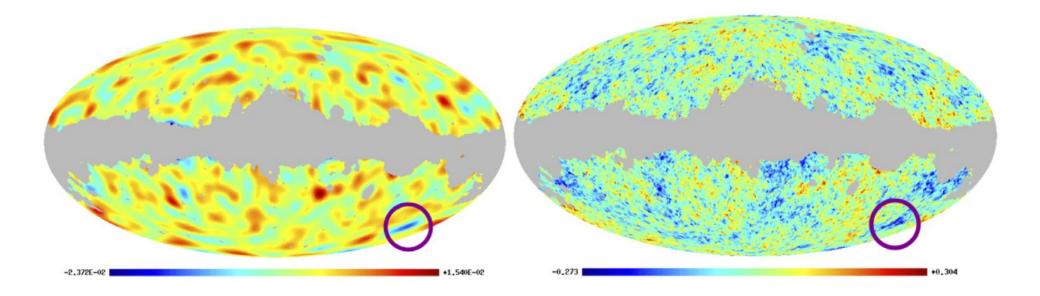
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### HEALPIX

- Hierarchical, Equal Area, isoLatitude
- Systematic effects?
- Few pixels per resolution element
- ~10<sup>6</sup> pixels x o(100) channels
- Objective:
  - Fourier analysis with spherical harmonics
  - Wavelet decomposition
  - Near neighbor searches



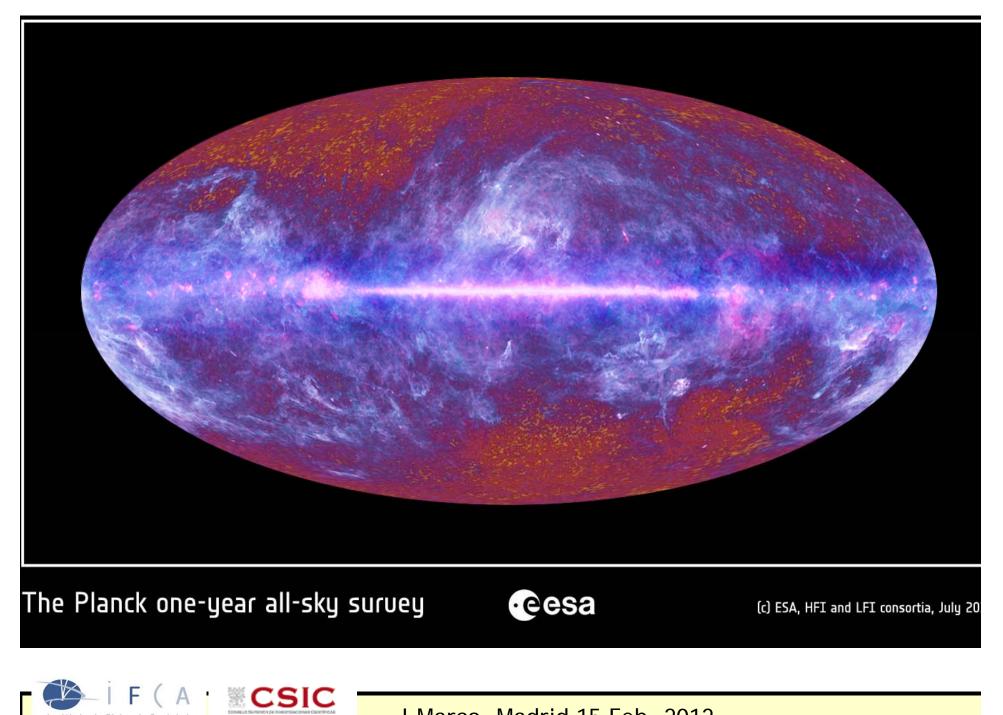
# Analysis: non-gaussianity sources



- Example of analysis on WMAP (P.Vielva, E.Martinez, et al. 2004)
- Skewness and the kurtosis of the SMHW (spherical Mexican hat wavelet) coefficients are calculated at different scales (ranging from a few arc minutes to tens of degrees).
- A non-Gaussian signal is detected at scales of the SMHW around 4<sup>o</sup> (size in the sky of around 10<sup>o</sup>)

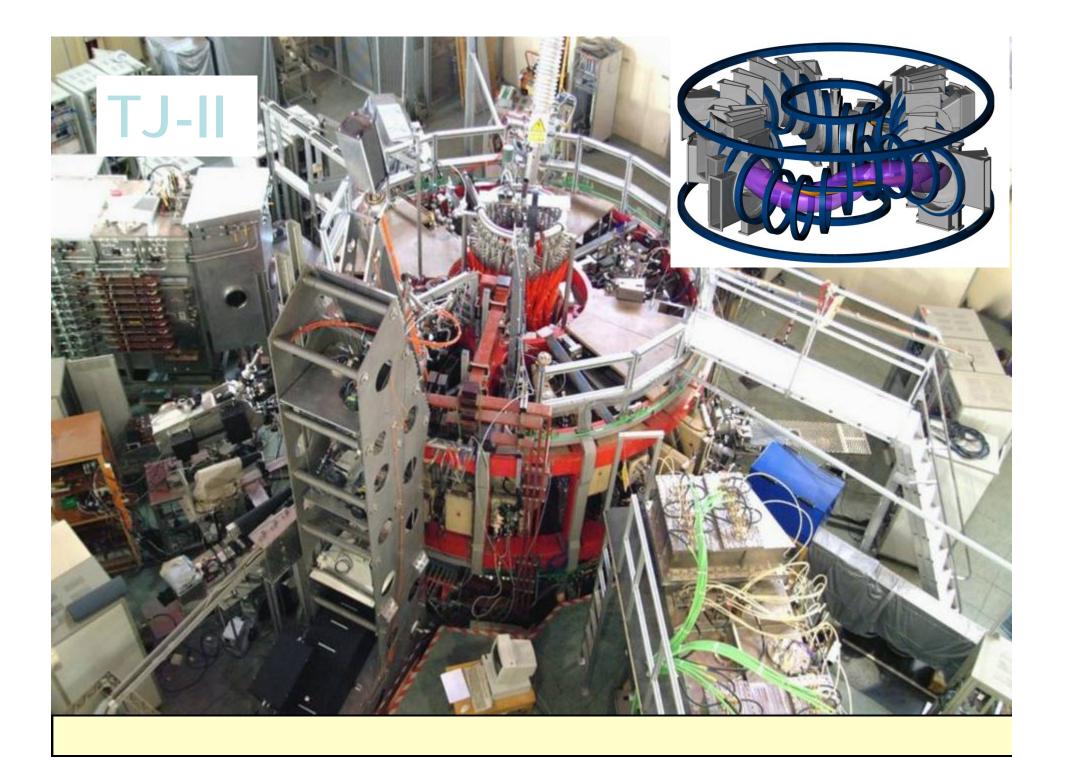




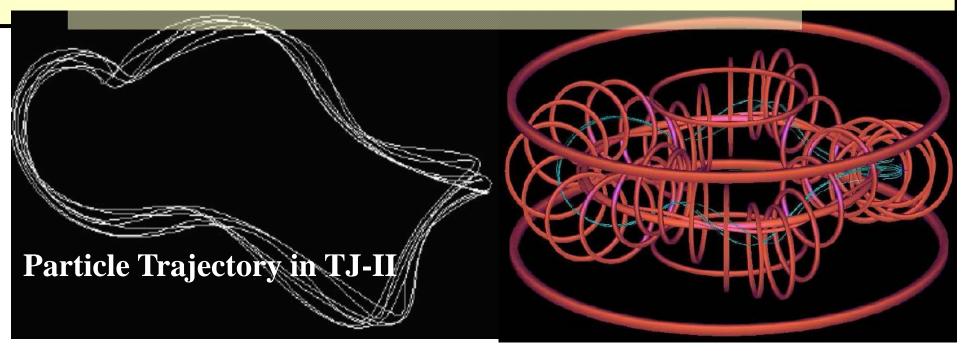


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### ISDEP: The Collisional Transport



-ISDEP solves the distribution function of a population of test particles in a fusion device: Test particles interact with a static plasma background.

**F. Castejon**, et al., Plasma Physics and Controlled Fusion 49 (2007) 753 J. L. Velasco, F. Castejón and A. Tarancón. Physics of Plasmas 16 (2009) 052303

### - Scales perfectly in distributed computing platforms

- Interactive European Grid: visualization

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### What else?

### Technology to handle large data volumes

- Large storage is not a large problem
  - (1 Petabyte ~ 300 K€ + 10KW 3TB disks, 4 GPFS servers, 8G FC+10GbEth)
- Fast processing "is" a problem:
  - Ex: SandyBridge processor + 2 SSD can handle up to 500MB/s
  - 100 GB in 1 node takes 200 s (base time)
  - 1 TB in 1 node will take 2000 s (base time)
  - While 10 TB in 10 nodes would take 200s
- Distributed or Parallel storage?
- Accommodate Interactive priority in the cluster/grid/cloud?
  - Dataflow oriented architecture





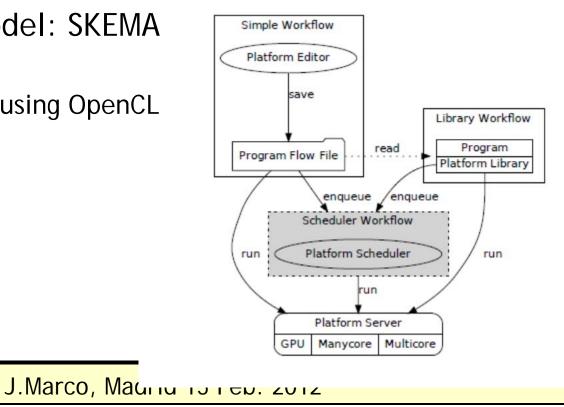
# What else?

- Data flow orientation + hardware integration
  - New hardware, GPUs in particular, is very powerful
  - (Efficiently) Programming new hardware is a problem
  - Handling dataflows is not an easy task
    - Example: EUFORIA project: KEPLER: supercomputing + grid
- Working on a new model: SKEMA
  - Graphical interface
  - Handling of hardware using OpenCL
  - First simple examples:
    - FFT

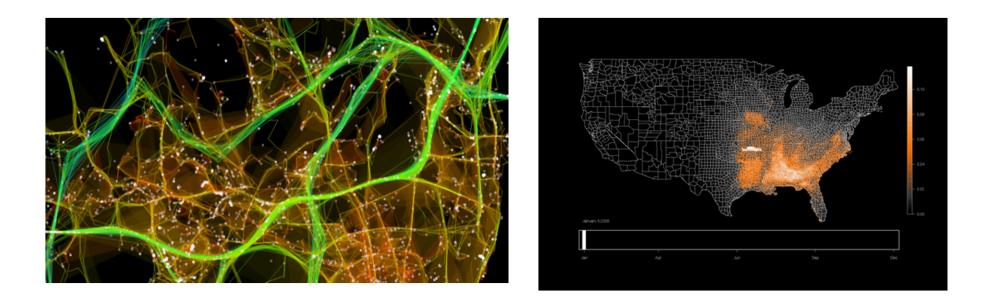
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Image Compression

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### What next?



### Global objective: pattern matching in a context

